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Massive vector fields on the Schwarzschild spacetime: quasi-normal modes and bound states

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Motivation

- Black hole perturbation theory
 - Massive vector fields less well understood
[Gal'tsov, Pomerantseva & Chizhov (1984); Konoplya (2006); Konoplya, Zhidenko & Molina (2007); Herdeiro, Sampaio & Wang (2011)]
 - Challenging problem (separability, coupled dof)
 - Rich perturbation spectrum
- Particle Phenomenology
 - Light U(1)'s in string theory [e.g. Jaeckel & Ringwald (2010)]
(RR-, NS-forms, D-branes w/ hyperweak couplings)
 - Kerr superradiance  Non-separable equations!
 - Schwarzschild case simpler and useful guide

Proca field on the Schwarzschild spacetime

Massive vector field equations

- Schwarzschild spacetime

$$ds^2 = -f(r)dt^2 + f^{-1}dr^2 + r^2(d\theta^2 + \sin^2\theta d\phi^2) , \quad f(r) = 1 - 2M/r$$

- Proca equations

$$F^{\mu\nu}_{;\mu} = \mu^2 A^\nu , \quad F_{\mu\nu} = A_{\nu;\mu} - A_{\mu;\nu}$$

- Lorenz condition:

$$A^\mu_{;\mu} = 0$$

- Field equations (Ricci-flat):

$$A^{\nu;\mu}_\mu - \mu^2 A^\nu = 0$$

Separation of variables

- Vector spherical harmonics [see e.g. Barack & Lousto (2005)]

$$A_\mu(t, r, \theta, \phi) = \frac{1}{r} \sum_{i=1}^4 \sum_{lm} c_i u_{(i)}^{lm}(t, r) Z_\mu^{(i)lm}(\theta, \phi)$$

$$\begin{aligned} Z_\mu^{(1)lm} &= [1, 0, 0, 0] Y^{lm} \\ Z_\mu^{(2)lm} &= [0, f^{-1}, 0, 0] Y^{lm} \\ Z_\mu^{(3)lm} &= \frac{r}{\sqrt{l(l+1)}} [0, 0, \partial_\theta, \partial_\phi] Y^{lm} \\ Z_\mu^{(4)lm} &= \frac{r}{\sqrt{l(l+1)}} \left[0, 0, \frac{1}{s_\theta} \partial_\phi, -s_\theta \partial_\theta \right] Y^{lm} \end{aligned}$$

Separation of variables

Using the Lorenz condition:

$$\begin{aligned}\hat{\mathcal{D}}_2 u_{(2)} & - \frac{2f}{r^2} \left(1 - \frac{3}{r}\right) (u_{(2)} - u_{(3)}) = 0 \\ \hat{\mathcal{D}}_2 u_{(3)} & + \left[\frac{2fl(l+1)}{r^2} u_{(2)} \right] = 0 \\ \hat{\mathcal{D}}_2 u_{(4)} & = 0\end{aligned}\quad \left. \begin{array}{l} \text{even-parity} \\ \text{odd-parity} \end{array} \right\}$$

where

$$\hat{\mathcal{D}}_2 \equiv -\frac{\partial^2}{\partial t^2} + \frac{\partial^2}{\partial r_*^2} - f \left[\frac{l(l+1)}{r^2} + \mu^2 \right]$$

Electromagnetic limit

- Monopole [Konoplya (2006)]

$$\left[-\frac{\partial^2}{\partial t^2} + \frac{\partial^2}{\partial r_*^2} - f \left(\frac{2(r-3)}{r^3} + \mu^2 \right) \right] u_{(2)} = 0$$

$$\downarrow \mu \rightarrow 0$$

$$\left[-\frac{\partial^2}{\partial t^2} + \frac{\partial^2}{\partial r_*^2} - \frac{2f}{r^3} \right] \dot{u}_{(1)} = 0$$

$$A_\mu = \chi_{,\mu} + \frac{q}{r} \delta_\mu^0$$

No monopole
radiation

- Odd-parity modes [Price (1972)]

$$\hat{\mathcal{D}}_2(r^2 \phi_1) = 0$$

where

$$\phi_1 \equiv F_{\mu\nu}(l^\mu n^\nu + \bar{m}^\mu m^\nu)$$

$$\phi_1^{lm} = i \frac{l(l+1)}{r^2} u_{(4)}^{lm}(t, r) Y_{lm}(\theta, \phi)$$

Electromagnetic limit

- Even-parity modes

$$\frac{1}{rf^3} \hat{\mathcal{D}}_{RW}^{l,s=1} \left[f^{-1} \hat{\mathcal{D}}_{RW}^{l,s=0} (ru_{(3)}) \right] = 0$$

$$\hat{\mathcal{D}}_{RW}^{l,s} \equiv -\frac{\partial^2}{\partial t^2} + \frac{\partial^2}{\partial r_*^2} - f \left(\frac{l(l+1)}{r^2} + \frac{2(1-s^2)}{r^3} \right)$$

Scalar solution

$$\hat{\mathcal{D}}_{RW}^{l,s=0} (ru_{(3)}) = 0$$

$$A_\mu = \chi_{,\mu} , \quad \chi = \frac{u_{(3)}}{l(l+1)} Y_{lm}(\theta, \phi)$$

(unphysical gauge mode)

Vector solution

$$\hat{\mathcal{D}}_{RW}^{l,s=1} \psi = \hat{\mathcal{D}}_2 \psi = 0$$

$$\psi = u'_{(3)} - \frac{l(l+1)}{r} u_{(2)}$$

(same for odd-parity modes)

Numerical methods

Frequency spectra

- Frequency-domain representation:

$$u_{(i)}^{lm}(t, r) = u_{(i)}^{lm}(\omega, r) e^{-i\omega t}$$

- ‘Ingoing’ boundary conditions: $u_{(i)}(\omega, r) \sim e^{-i\omega r_*}$

- Asymptotic solution:

$$u_{(i)}(\omega, r) \sim B_{(i)}(\omega) e^{-qr_*} + C_{(i)}(\omega) e^{qr_*}$$



$$q = \sqrt{\mu^2 - \omega^2}$$

(quasi-)
bound states

quasi-normal
modes

Continued-fraction

- Ansatz: [Leaver (1985), see also Dolan (2007)]

$$u_{(i)}(\omega, r) = f^{-2i\omega} r^{-\nu} e^{qr} \sum_n a_n^{(i)} [f(r)]^n \quad \nu = (\omega^2 - q^2)/q$$

- Three-term recurrence relations:

$$\begin{aligned} \alpha_0 a_1 + \beta_0 a_0 &= 0 \\ \alpha_n a_{n+1} + \beta_n a_n + \gamma_n a_{n-1} &= 0, \quad n > 0 \end{aligned}$$



$$\beta_n - \frac{\alpha_{n-1} \gamma_n}{\beta_{n-1} - \frac{\alpha_{n-2} \gamma_{n-1}}{\beta_{n-2} - \alpha_{n-3} \gamma_{n-2} / \dots}} = \frac{\alpha_n \gamma_{n+1}}{\beta_{n+1} - \frac{\alpha_{n+1} \gamma_{n+2}}{\beta_{n+2} - \alpha_{n+2} \gamma_{n+3} / \dots}}$$

(monopole and odd-parity modes)

Continued-fraction

- Even-parity modes

$$\begin{aligned} \underline{\alpha}_0 \mathbf{U}_1 + \underline{\beta}_0 \mathbf{U}_0 &= 0, \\ \underline{\alpha}_n \mathbf{U}_{n+1} + \underline{\beta}_n \mathbf{U}_n + \underline{\gamma}_n \mathbf{U}_{n-1} &= 0, \quad n > 0 \end{aligned} \quad \mathbf{U}_n = \begin{pmatrix} a_n^{(2)} \\ a_n^{(3)} \end{pmatrix}$$

- Matrix continued fraction [Simmendinger *et al.* (1999)]

$$\mathbf{U}_{n+1} = \mathbf{R}_n^+ \mathbf{U}_n \quad \mathbf{R}_n^+ = - \left(\underline{\beta}_{n+1} + \underline{\alpha}_{n+1} \mathbf{R}_{n+1}^+ \right)^{-1} \underline{\gamma}_{n+1}$$

$$\mathbf{M} \equiv \underline{\beta}_0 - \underline{\alpha}_0 \left[\underline{\beta}_1 - \underline{\alpha}_1 \left(\underline{\beta}_2 + \underline{\alpha}_2 \mathbf{R}_2^+ \right) \underline{\gamma}_2 \right]^{-1} \underline{\gamma}_1$$

$$\mathbf{M} \mathbf{U}_0 = 0 \quad \rightarrow \quad \det |\mathbf{M}| = 0$$

- “Polarization”:

$$\mathcal{P} \equiv \lim_{r_* \rightarrow \infty} \frac{u_{(3)}(\omega, r)}{u_{(2)}(\omega, r)} = \sum_{n=0}^{\infty} a_n^{(3)} / \sum_{n=0}^{\infty} a_n^{(2)}$$

Forward-integration (bound states)

- “Ingoing” *ansatz*:

$$u_{(i)}(\omega, r) = (r - r_H)^{-2i\omega} \sum_{n=0}^{\infty} b_{(i)n}(r - r_H)^n$$

- Bound state solutions converge exponentially
- Integrate eqs. up to large distance and minimize in complex frequency-plane
- For **even-parity** modes, choose basis of coefficients and integrate:

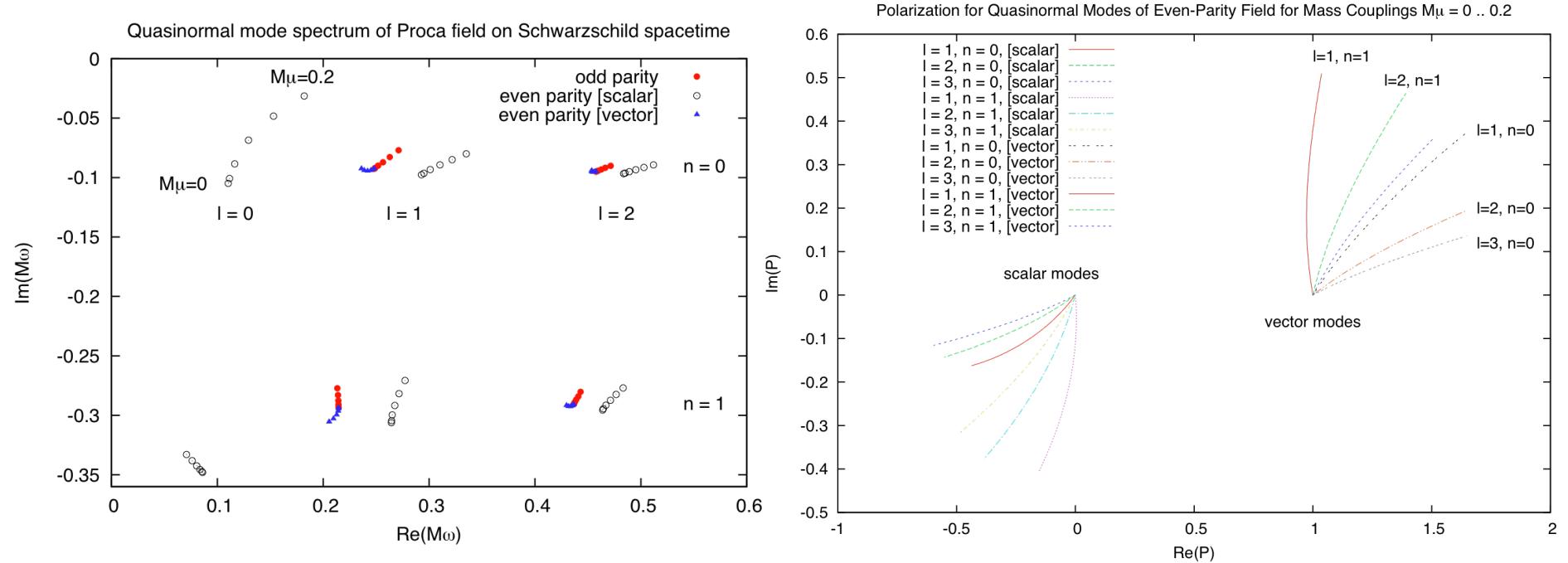
$$\mathbf{S}(\omega, r) = \begin{pmatrix} u_{(2)}^{(1,0)}(\omega, r) & u_{(2)}^{(0,1)}(\omega, r) \\ u_{(3)}^{(1,0)}(\omega, r) & u_{(3)}^{(0,1)}(\omega, r) \end{pmatrix}$$

- Bound states minimize

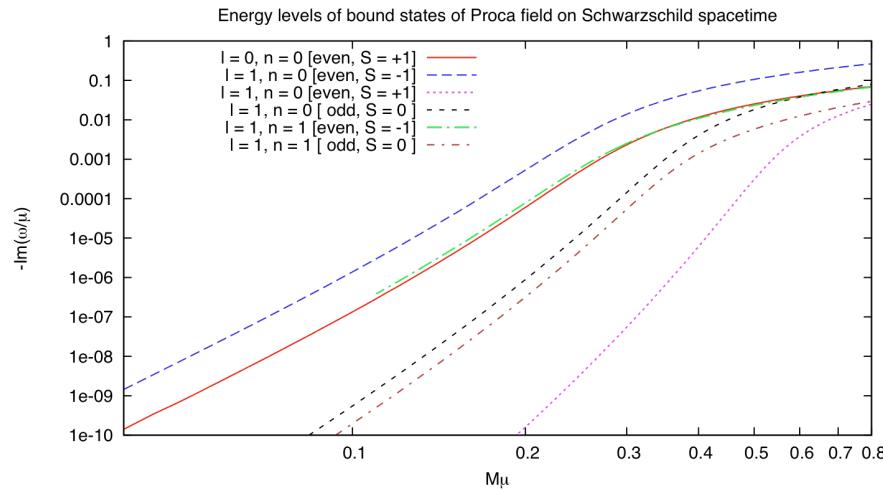
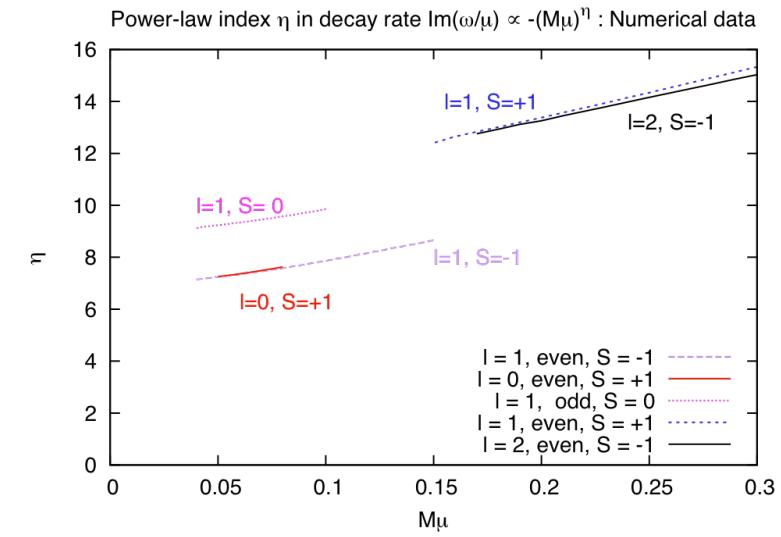
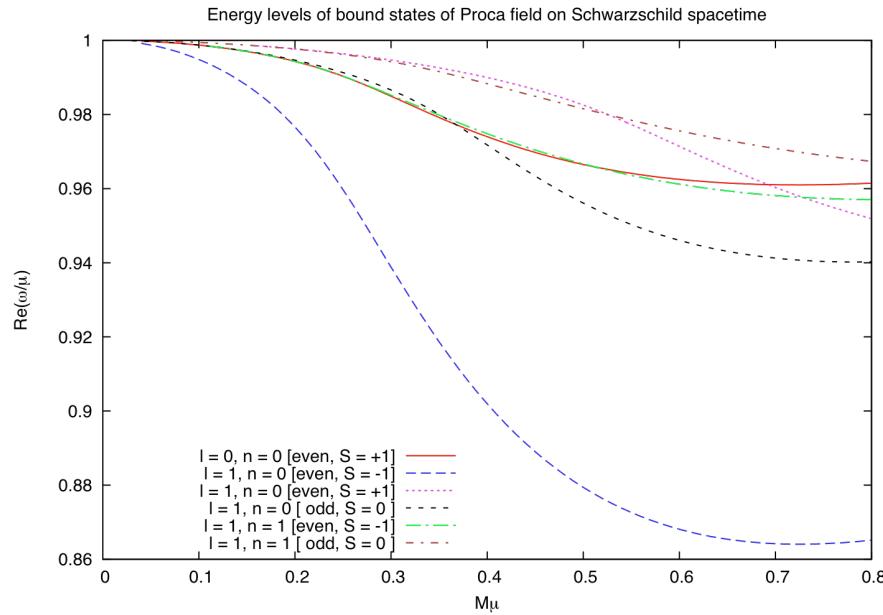
$$\det |\mathbf{S}(\omega, r \rightarrow +\infty)|$$

Results

Quasi-normal modes



Bound states: frequency spectrum



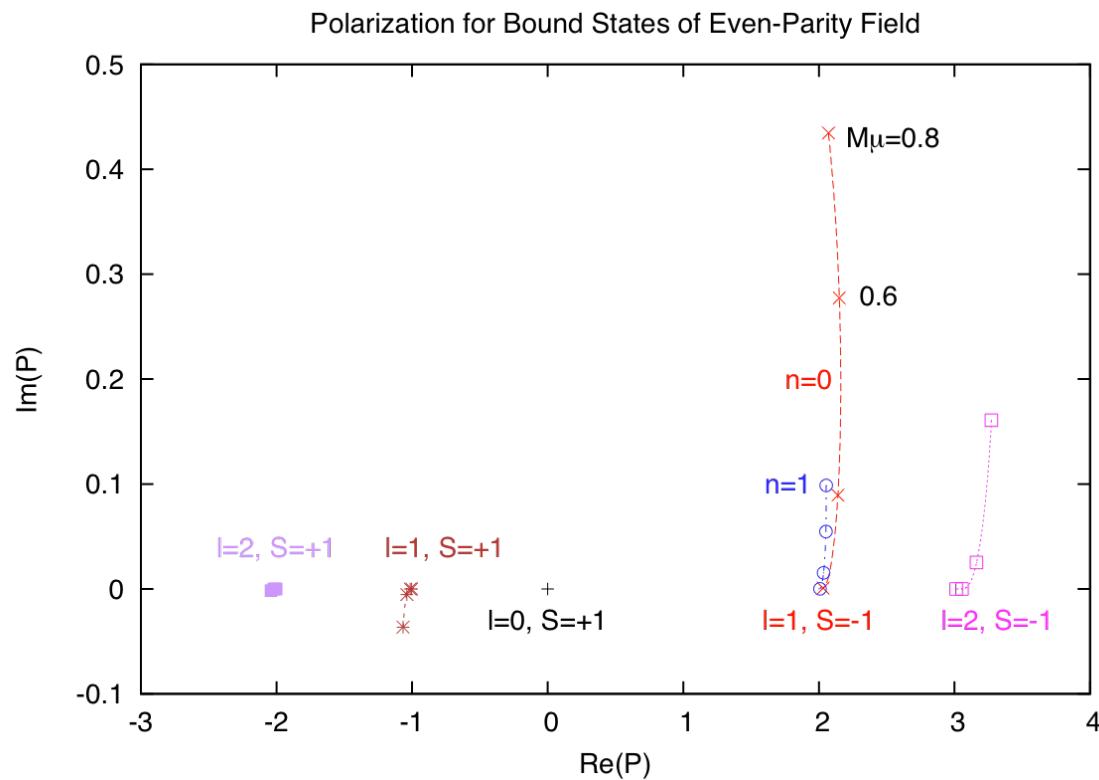
$$\text{Re}(\omega/\mu) \approx 1 - \frac{(M\mu)^2}{2(l + S + 1 + n)^2}$$

[Gal'tsov, Pomerantseva & Chizhov (1984)]

$$\text{Im}(\omega/\mu) \approx (M\mu)^\eta$$

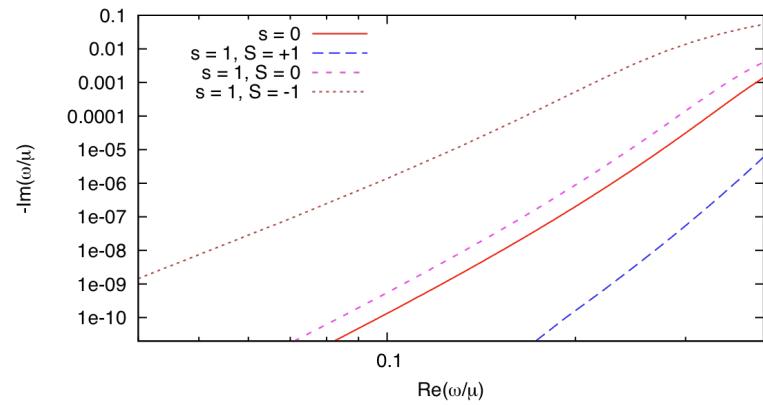
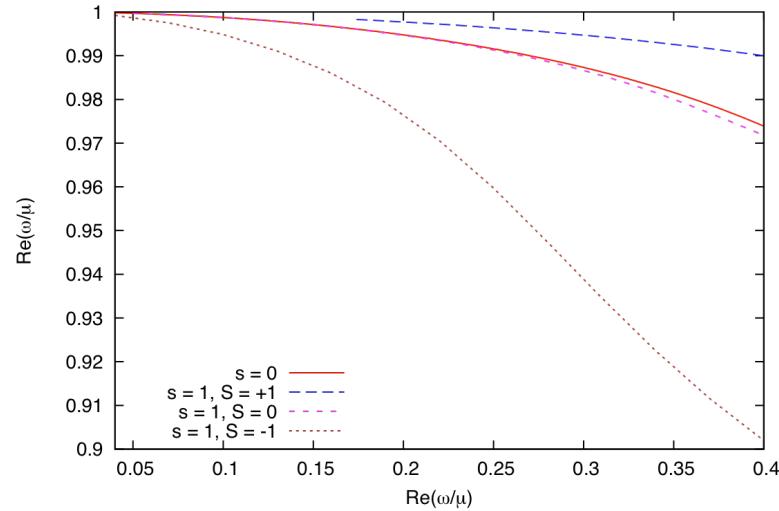
$$\eta = 4l + 2S + 5$$

Bound states: polarization

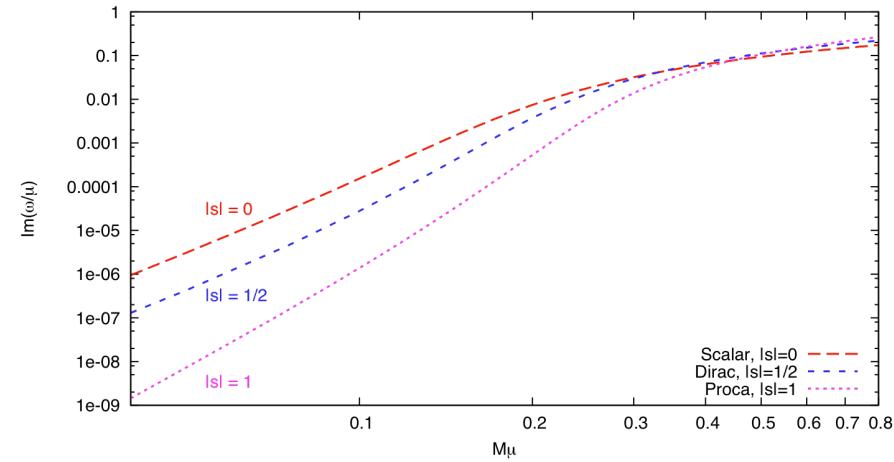
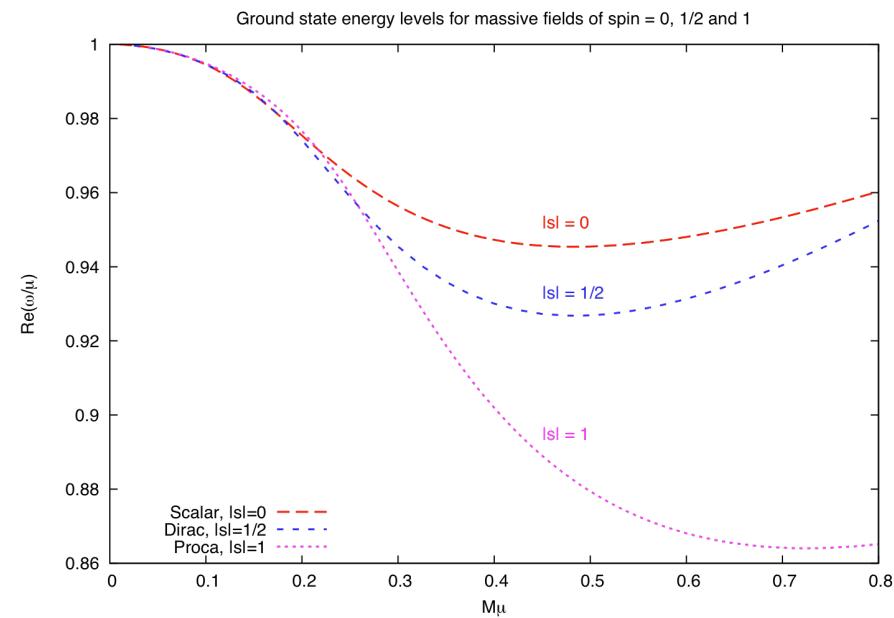


$$\mathcal{P} \rightarrow -l, l+1, \quad S = \pm 1$$

Bound states: comparison with other spins



$|l=1, n=0$ modes



Analytical results (small mass coupling)

Odd-parity modes

$$[x^2(x+1)^2\partial_x^2 + x(x+1)\partial_x + V(x)]u_{(4)} = 0 \quad x = r/2M - 1$$

$$V(x) = 4\omega^2(x+1)^4 - 4\mu^2x(x+1)^3 - l(l+1)x(x+1)$$

- **Near region:** $\mu x, \omega x \ll l$

$$u_{(4)}^{near} = A_{(4)}x^{-2i\omega}(x+1)^{1+\delta} {}_2F_1(-l-2i\omega+\delta, l+1-2i\omega+\delta, 1-4i\omega, -x)$$

$$\delta = \sqrt{1 - 4\omega^2}$$

- **Far region:** $x \gg 1$

$$u_{(4)}^{far} = e^{-z/2} [C_{(4)}z^{l+1}M(l+1-\nu, 2l+2, z) + D_{(4)}z^{-l}M(-l-\nu, -2l, z)]$$

$$z = 4qx$$

Odd-parity modes

- Matching for bound-states: $\omega \simeq \mu \left(1 - \frac{\mu^2}{2(l+1+n)^2} \right)$

$$\delta\omega \simeq \frac{4^{2l+1}\mu^{4l+5}}{(l+1+n)^{2l+4}} \frac{(2l+1+n)!}{n!} \frac{(l+1)!(l-1)!}{[(2l)!(2l+1)!]^2} (1 + 2i\omega) \prod_{k=1}^l (k^2 - 1 - 4i\omega)$$

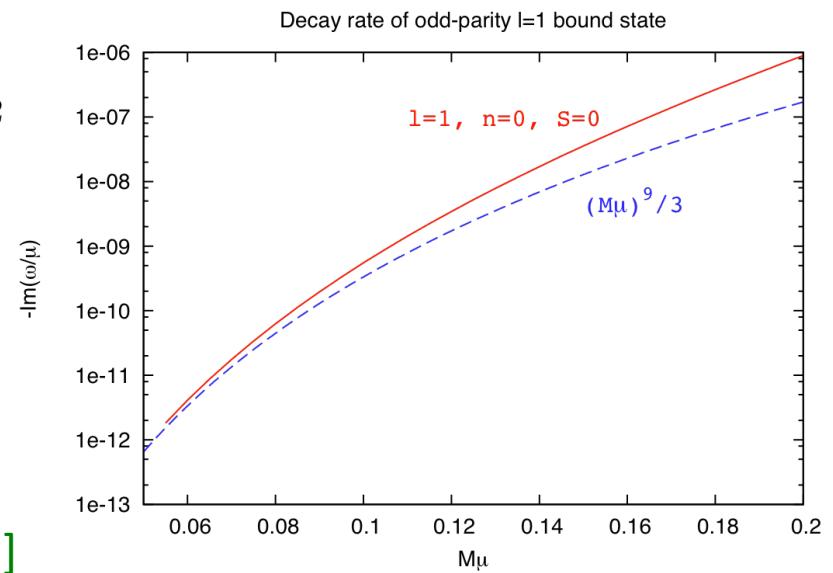
- Matching for QN modes:

$$1 = i(4q)^{2l+1}(l+1)!(l-1)! \left(\frac{l!}{(2l)!(2l+1)!} \right)^2 \times (1 + 2i\omega) \prod_{k=1}^l (k^2 - 1 - 4i\omega)$$

- Bad approximation for QN modes

[see also Panchapakesan & Majumdar (1987)]

- Correct mass dependence



Even-parity modes

$$\begin{aligned}[x^2(x+1)^2\partial_x^2 + x(x+1)\partial_x + V(x)]u_{(2)} - x(2x-1)(u_{(2)} - u_{(3)}) &= 0 \\ [x^2(x+1)^2\partial_x^2 + x(x+1)\partial_x + V(x)]u_{(3)} + 2\lambda^2 x(x+1)u_{(2)} &= 0\end{aligned}$$

- Diagonalization in far region:

$$u_{(2,3)}(z) = \sum_{S=\pm 1} c_{(2,3)}^S u_{(S)}(z)$$

$$u_{(S)}^{far} = e^{-z/2} [C_{(S)} z^{j+1} M(j+1-\nu, 2j+2, z) + D_{(S)} z^{-j} M(-j-\nu, -2j, z)]$$

- No decoupling in near region...

$$\begin{aligned}[x^2(x+1)^2\partial_x^2 + x(x+1)\partial_x + V(x)]\psi &= 4\mu^2 x(x+1)u_{(3)} , \\ [x^2(x+1)^2\partial_x^2 + x(x+1)(2x+1)\partial_x + V(x)]u_{(3)} &= 2x(x+1)^2\psi\end{aligned}$$

Even-parity modes

$$\begin{aligned}[x^2(x+1)^2\partial_x^2 + x(x+1)\partial_x + V(x)]u_{(2)} - x(2x-1)(u_{(2)} - u_{(3)}) &= 0 \\ [x^2(x+1)^2\partial_x^2 + x(x+1)\partial_x + V(x)]u_{(3)} + 2\lambda^2 x(x+1)u_{(2)} &= 0\end{aligned}$$

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- No decoupling in near region...

$$\begin{aligned}[x^2(x+1)^2\partial_x^2 + x(x+1)\partial_x + V(x)]\psi &= \cancel{4\mu^2} x(x+1)u_{(3)} , \\ [x^2(x+1)^2\partial_x^2 + x(x+1)(2x+1)\partial_x + V(x)]u_{(3)} &= 2x(x+1)^2\psi\end{aligned}$$

Even-parity modes

What can we conclude?

- Hydrogenic spectrum for bound states from far region solutions
$$\omega \simeq \mu \left(1 - \frac{\mu^2}{2(j+1+n)^2} \right), \quad j = l \pm 1$$
- Asymptotic solutions are linear combinations of ‘scalar’ and ‘vector’ solutions for QN modes
- ‘scalar-vector’ mixing in far region yields non-trivial spin-dependence for imaginary bound state frequency:

$$\text{Im}(\omega/\mu) \approx (M\mu)^\eta \quad \eta = 4l + 2S + 5$$

Summary and future prospects

- Proca fields have a rich spectrum of perturbations
- Decoupling of odd- and even-parity modes
- Broken degeneracy for ‘vector’ QN modes + physical ‘scalar’ mode
- Hydrogen-like bound states with spin-dependent decay rates
- Larger decay rates for $S=-1,0$ modes than for scalar fields



- Larger superradiant instabilities in Kerr spacetimes?
- Encouraging preliminary results for slowly-rotating BH
- Interesting phenomenology for light string U(1) fields!