

VII Black Holes Workshop

Aveiro, Portugal, December, 18-19, 2014

Non-minimal black holes with regular electric field

Alexander Balakin

Kazan Federal University, Institute of Physics

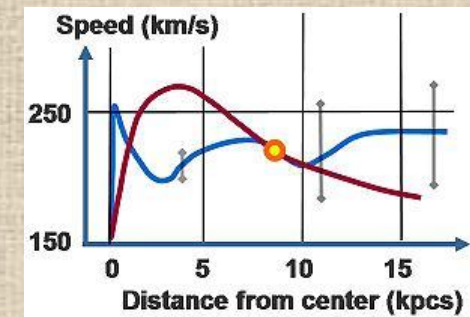


Plan of the talk

- Motivation and problems
- Mathematical formalism
- Examples of exact solutions
- Summary

Basic motives for the non-minimal extensions of the Field Theory

- **Non-minimal interactions as an *alternative to Dark Energy*:** is it possible to explain accelerated expansion of the Universe in terms of curvature coupling?
- **Non-minimal interactions as an *alternative to Dark Matter*:** is it possible to explain flat rotation curves of spiral galaxies in terms of curvature coupling? (Navarro–Frenk–White)
- **Non-minimal interactions and causal structure of space-time:** can non-minimal interactions eliminate or transform *singularities and horizons*?



$$U = \frac{GM}{r} \quad ??? \quad \Rightarrow \quad ??? \quad U^* = \frac{GM}{\sqrt{r^2 + a^2}}$$

Bardeen, 1975

$$U^*(0) = \frac{GM}{a}$$

Our dream is to find

exact static spherically symmetric solutions to the master equations of the non-minimal Einstein-Maxwell theory (or its extension), which satisfy the following three conditions:

1. The solutions for electromagnetic and gravitational fields have to be **analytical** (explicit or in quadratures) .
2. The solutions for magnetic and electric fields have to be **regular in the center**.
3. The solutions for the metric functions and the corresponding curvature invariants have to be **regular in the center**.

The main *problem* of the Non-minimal Field Theory seems to be connected with a large number of coupling constants introduced phenomenologically:

- 1. NM Einstein-Maxwell theory contains 3 new coupling parameters.**
- 2. NM Einstein-Yang-Mills-Higgs theory includes 8 new parameters.**

How one can reduce the number of parameters, e.g., to one (non-minimal radius)?

- A) To use the requirements of *regularity* of the metric !**
- B) To reduce new coupling parameters to *known* constants!**
- C) To introduce geometric analogs for NM *susceptibilities* !**

Non-minimal Einstein-Maxwell model

$$S_{\text{NMEM}} = \int d^4x \sqrt{-g} \left[\frac{R}{8\pi} + \frac{1}{2} F_{ik} F^{ik} + \frac{1}{2} \mathcal{R}^{ikmn} F_{ik} F_{mn} \right]$$

Non-minimal susceptibility tensor

$$\mathcal{R}^{ikmn} = q_1 R g^{ikmn} + q_2 \mathfrak{R}^{ikmn} + q_3 R^{ikmn}$$

$$g^{ikmn} \equiv \frac{1}{2} (g^{im} g^{kn} - g^{in} g^{km}),$$

$$\mathfrak{R}^{ikmn} \equiv \frac{1}{2} (R^{im} g^{kn} - R^{in} g^{km} + R^{kn} g^{im} - R^{km} g^{in})$$

Non-minimally extended Maxwell equations

$$\nabla_k H^{ik} = I^i$$

$$H^{ik} \equiv F^{ik} + \mathcal{R}^{ikmn} F_{mn}$$

$$\nabla_k F^{*ik} = 0$$

Three examples of relations between non-minimal coupling constants

I. Drummond-Hathrell -type relations, based on QED one-loop calculations:

$$q_1 \equiv -5\tilde{Q}, q_2 = 13\tilde{Q}, q_3 = -2\tilde{Q}, \quad \tilde{Q} \equiv \frac{\alpha \lambda_e^2}{180\pi}$$

Fine structure constant

Compton wave-length of electron

II. Gauss-Bonnet type requirements: NM susceptibility tensor is proportional to the double dual Riemann tensor

$$q_1 + q_2 + q_3 = 0 \quad + \quad 2q_1 + q_2 = 0 \quad \longrightarrow \quad \mathcal{R}_{ikmn} = \gamma^* R_{ikmn}^*$$

III. NM susceptibility tensor is proportional to the difference between Riemann and Weyl tensors

$$3q_1 + q_2 = 0 \quad + \quad q_3 = 0 \quad \longrightarrow \quad \mathcal{R}_{ikmn} = \Omega [R_{ikmn} - \mathcal{C}_{ikmn}]$$

Static spherically symmetric models with electric field

Balakin A.B., Bochkarev V.V. and Lemos J.P.S. Phys. Rev. D, 2008.

Balakin A.B., Lemos J.P.S. and Zayats A.E. Phys. Rev. D, 2010

$$ds^2 = \sigma^2 N dt^2 - \frac{dr^2}{N} - r^2(d\theta^2 + \sin^2 \theta d\varphi^2)$$

$$\sigma(\infty) = 1$$

$$N(\infty) = 1$$

Equation for the electric field is coupled to metric functions and their derivatives

$$E(r) \left\{ r^2 \left[1 + (q_1 + q_2 + q_3) \left(N'' + 3N' \frac{\sigma'}{\sigma} + 2N \frac{\sigma''}{\sigma} \right) \right] \right. \\ \left. + 2r(2q_1 + q_2) \left(N' + N \frac{\sigma'}{\sigma} \right) + 2q_1(N - 1) \right\} = Q$$

$$E^2(r) = -\frac{1}{2} F_{ik} F^{ik}$$

$$\begin{aligned}
\frac{[r(1-N)]'}{\kappa r^2} = & -(E^2)''N(q_1 + q_2 + q_3) \\
& + (E^2)' \left[-\frac{1}{2}(q_1 + q_2 + q_3) \left(N' + \frac{8N}{r} \right) \right. \\
& \left. + \frac{N}{r}(2q_1 + q_2) \right] + E^2 \left[\frac{1}{2} + (q_1 + q_2 + q_3) \right. \\
& \times \left(N'' + 3N' \frac{\sigma'}{\sigma} + 2N \frac{\sigma''}{\sigma} - \frac{N'}{r} - 2 \frac{N}{r^2} \right) \\
& + (2q_1 + q_2) \left(2 \frac{N'}{r} + 2 \frac{N}{r} \frac{\sigma'}{\sigma} + \frac{N}{r^2} \right) \\
& \left. + q_1 \frac{(N-1)}{r^2} \right], \tag{16}
\end{aligned}$$

**Two master equations
for the gravity field**

*(we write them simply to demonstrate
the scale of problems with integration
procedure for arbitrary set of coupling
parameters)*

$$\begin{aligned}
\frac{2\sigma'}{\kappa r \sigma} = & -(E^2)''(q_1 + q_2 + q_3) \\
& + (E^2)' \left[(q_1 + q_2 + q_3) \left(\frac{\sigma'}{\sigma} - \frac{4}{r} \right) + (2q_1 + q_2) \frac{2}{r} \right] \\
& + E^2 \left[(q_1 + q_2 + q_3) \frac{2\sigma'}{r\sigma} - \frac{2q_3}{r^2} \right]. \tag{17}
\end{aligned}$$

Guiding parameters of the model and convenient dimensionless quantities

Reissner-Nordstrom radius

$$r_Q \equiv \sqrt{G}|Q|$$

Schwarzschild radius

$$r_M \equiv 2GM$$

Non-minimal radius

$$r_q = \sqrt{2|q|}$$

$$a \equiv \frac{2q}{r_Q^2} = \pm \frac{r_q^2}{r_Q^2}$$

Dimensionless
guiding parameter

$$x = \frac{r}{r_Q}$$

$$E_Q \equiv \frac{Q}{r_Q^2}$$

$$Z(x) = \frac{E(r)}{E_Q}$$

Dimensionless
electric field

I. Exactly integrable model with

$$q_1 = -q_2, \quad q_3 = 0$$

!!! Special case, when the non-minimal and Reissner-Nordstrom's radii coincide

$$a = 1$$

Equation for the electric field transforms into cubic algebraic equation:

$$(Z - 1)[(1 + x^2)Z^2 + (x^2 - 1)Z - 1] = 0$$

$$Z_{\text{Coulomb}}(x, 1) = \frac{1}{2(1 + x^2)} [1 - x^2 + \sqrt{x^4 + 2x^2 + 5}]$$

$$Z(x, 1) \rightarrow \frac{1}{x^2} \quad E(r) \rightarrow \frac{Q}{r^2}$$

**Regular
electric field**

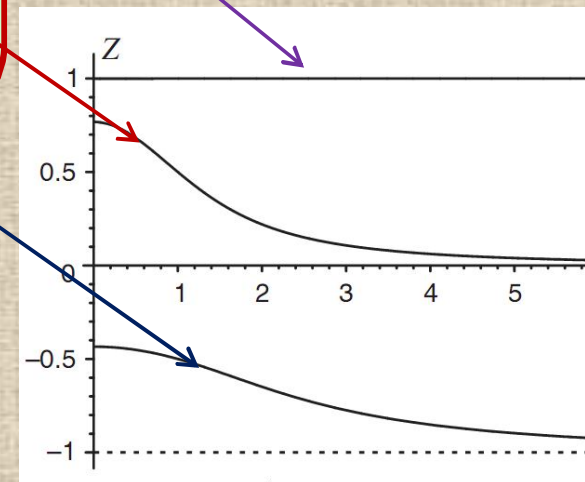
$$Z(0, 1) = \frac{\sqrt{5}+1}{2} = 1.618\dots \quad \text{"golden section"} \quad \phi$$

$$Z_{\text{const}}(x, 1) = 1 \quad E(r) = E_Q$$

$$Z_{\text{nonCoulomb}}(x, 1) = \frac{1}{2(1+x^2)} [1 - x^2 - \sqrt{x^4 + 2x^2 + 5}]$$

$$Z_{\text{nonCoulomb}}(0, 1) = -\frac{2}{1 + \sqrt{5}} = -\frac{1}{\phi}$$

$$Z_{\text{nonCoulomb}}(r \rightarrow \infty, 1) \rightarrow -1 - \frac{1}{x^4}$$



I. Exactly integrable model with Regular metric functions

$$q_1 = -q_2, \quad q_3 = 0$$

$$\sigma(x) = \exp\left\{-\frac{3 + (1 - x^2)\sqrt{x^4 + 2x^2 + 5} + x^4}{2(1 + x^2)^2}\right\}$$

$$\sigma(\infty) = 1$$

$$\sigma(0) = \exp\left\{-\frac{1}{2}(3 + \sqrt{5})\right\} \longrightarrow \exp\{-(1 + \phi)\}$$

$$N(x) = \frac{1}{2x\sigma(x)} \int_0^x d\xi \sigma(\xi) [\xi^2 + 3 - \sqrt{\xi^4 + 2\xi^2 + 5}]$$

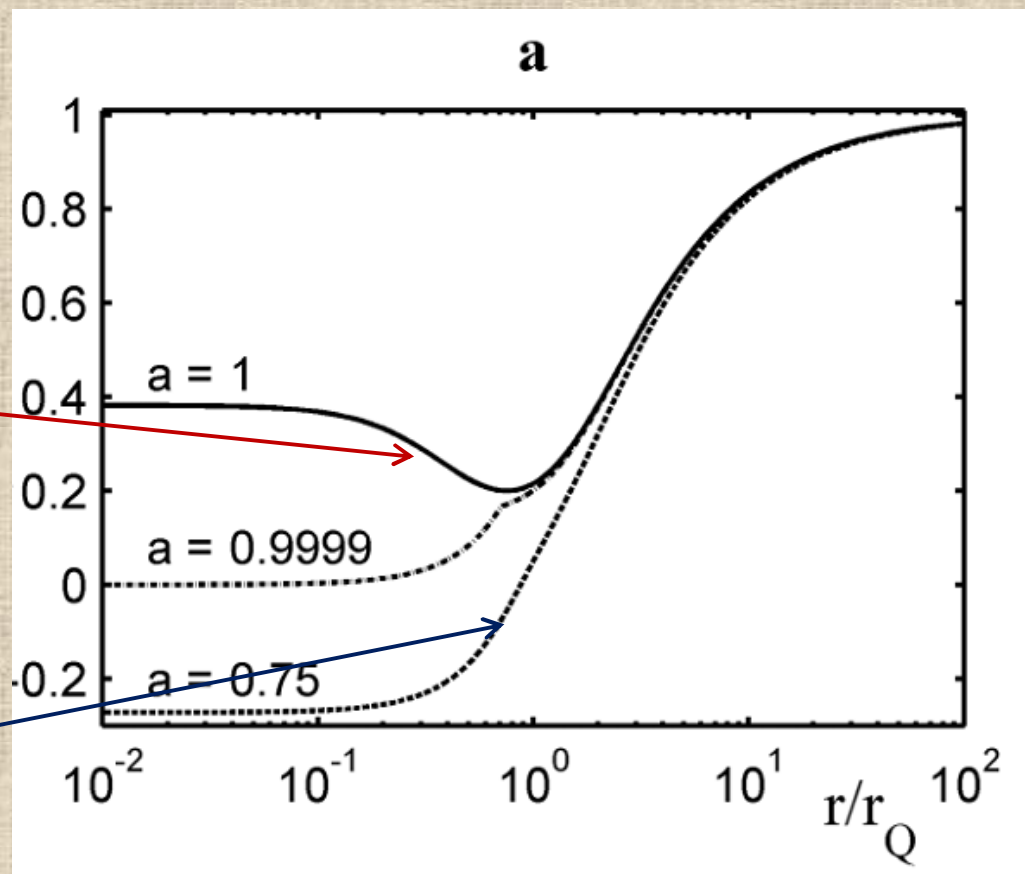
$$N(0) = \frac{3 - \sqrt{5}}{2}$$

$$N(\infty) = 1$$

**Metric function $N(r)$ is regular and positive everywhere.
There are no horizons, i.e., this exact solution does not describe black holes.
We deal with the so-called regular Fibonacci soliton...**

Fibonacci soliton

Example of
black hole



$N(r)$

Asymptotic mass

$$M = \frac{|Q|}{4\sqrt{G}} \int_0^\infty d\xi \left[\frac{1}{\sigma(\xi)} - \xi \frac{\sigma'(\xi)}{\sigma^2(\xi)} - \frac{1}{2} \sigma(\xi) (\xi^2 + 3 - \sqrt{\xi^4 + 2\xi^2 + 5}) \right]$$

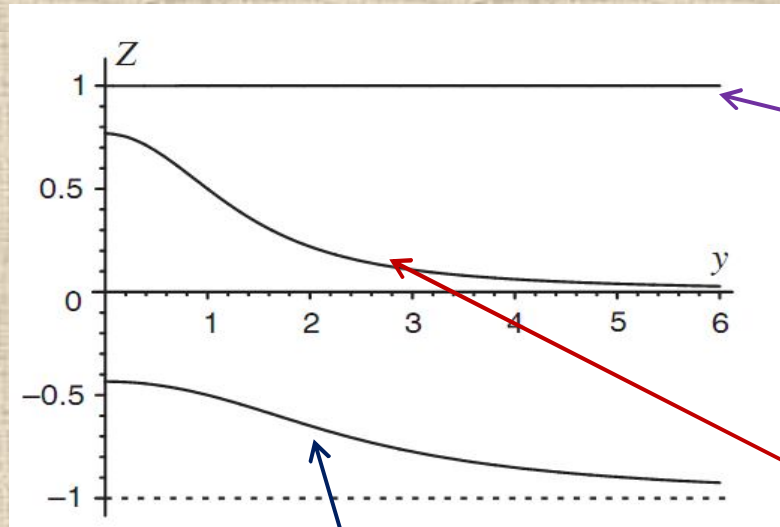
$$M \simeq 0.442 \frac{|Q|}{\sqrt{G}}$$

II. Exactly integrable model with $q_1 = -q, q_2 = 3q, q_3 = 0$

$$a = \frac{2q}{r_Q^2},$$

$$a = 1$$

Behavior of the electric field



$$Z = 1$$

Coulombian-type solution

$$Z_*(y, 1) = \frac{\sqrt{13 + 2y^2 + y^4} + 1 - y^2}{2(3 + y^2)}$$

Non-Coulombian-type solution

$$Z_-(y, 1) = \frac{-\sqrt{13 + 2y^2 + y^4} + 1 - y^2}{2(3 + y^2)}$$

$$Z_-(0, 1) = \frac{1 - \sqrt{13}}{6}$$

$$Z_*(0, 1) = \frac{1 + \sqrt{13}}{6}$$

Behavior of the metric function

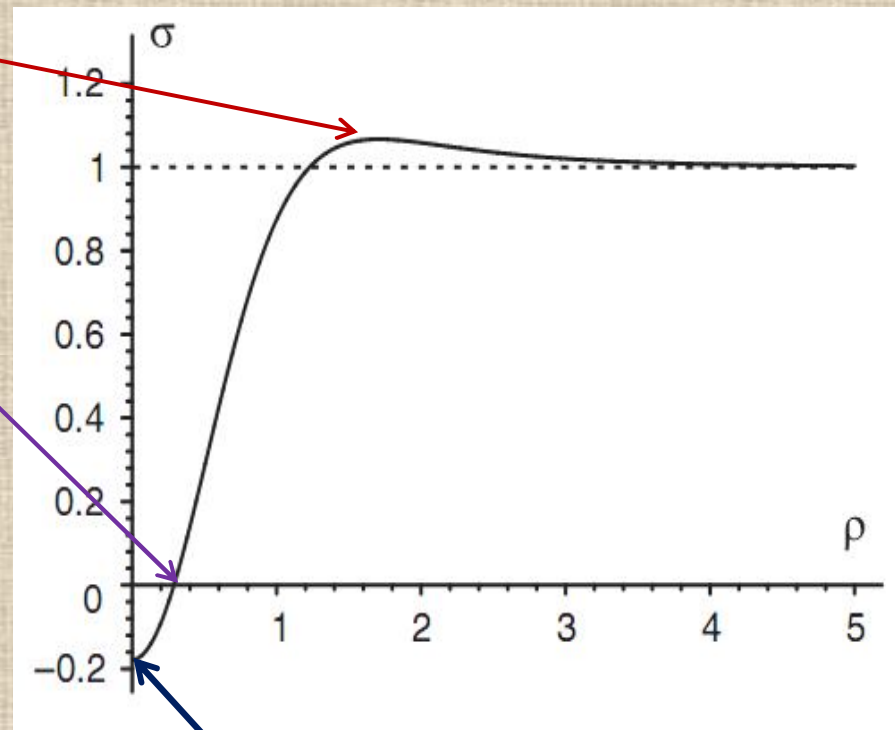
$$\sigma(\rho)$$

$$\sigma(\rho) = \frac{(6 + 34\rho^2)\sqrt{13 + 2\rho^2 + \rho^4} - 39 + 82\rho^2 + 10\rho^4 + 10\rho^6 + \rho^8}{(3 + \rho^2)^3\sqrt{13 + 2\rho^2 + \rho^4}}.$$

$$\sigma_{\max} = 1.067$$

$$\rho = 1.706$$

$$\rho_s = 0.289$$



$$\sigma(0, 1) = -(\sqrt{13} - 2)/9 \approx -0.178$$

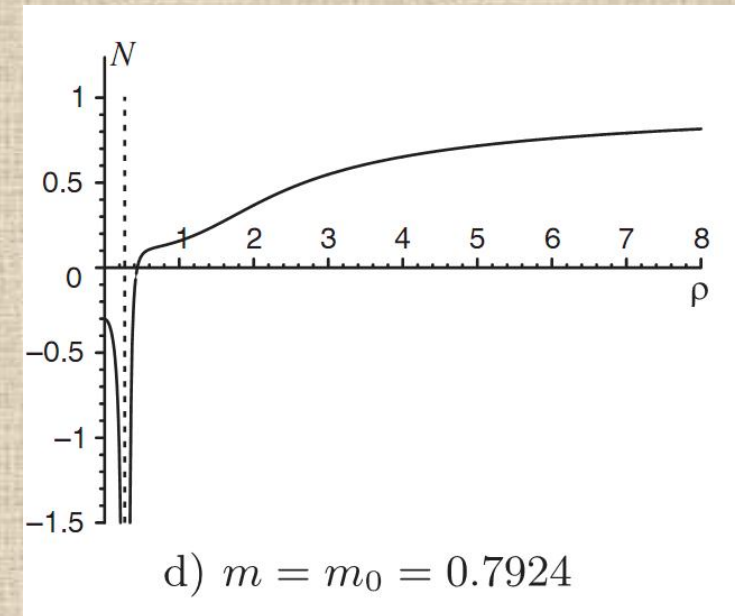
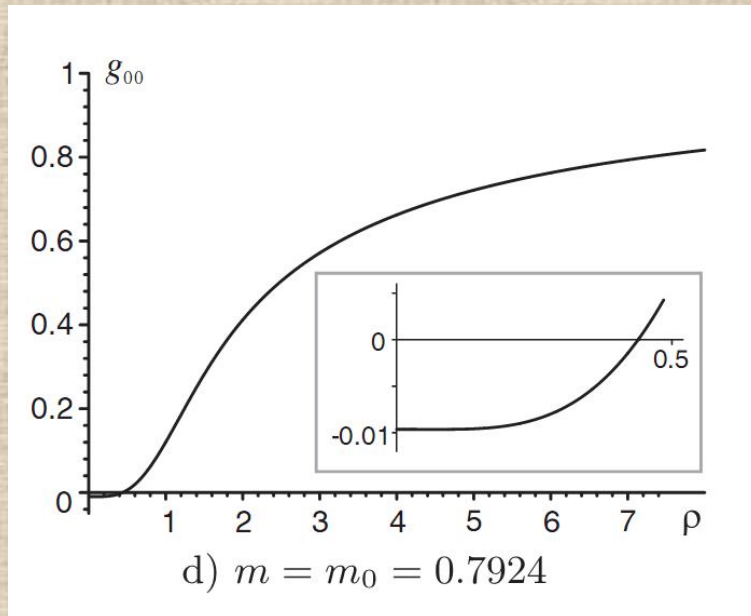
Behavior of the metric function $N(\rho)$

$$N = \frac{1}{\sigma^2} \left[1 - \frac{2m - J(\rho)}{\rho} \right]$$

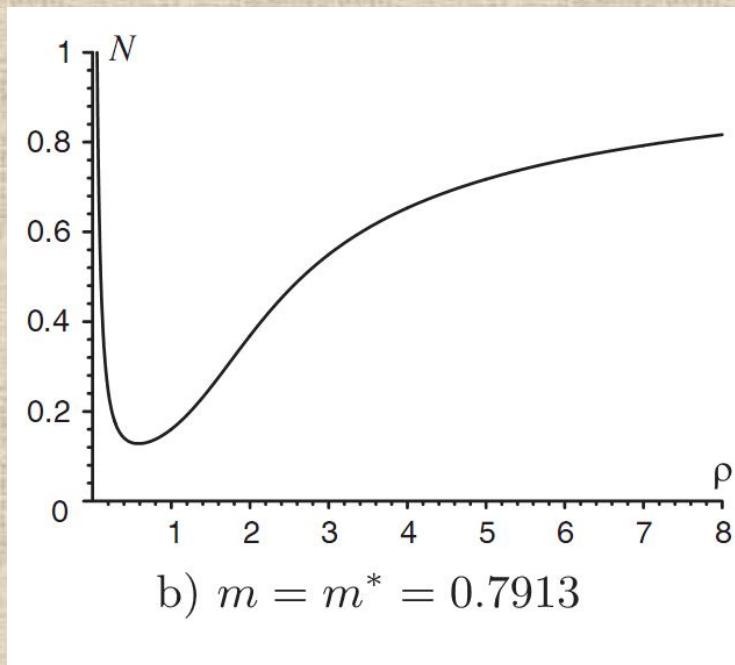
$$J(\rho) = \int_{\rho}^{\infty} d\rho \left[1 - \frac{\sigma}{2(3 + \rho^2)^2} (19 + 23\rho^2 + 5\rho^4 + \rho^6 - (5 + 2\rho^2 + \rho^4)\sqrt{13 + 2\rho^2 + \rho^4}) \right]$$

First specific value of the asymptotic mass

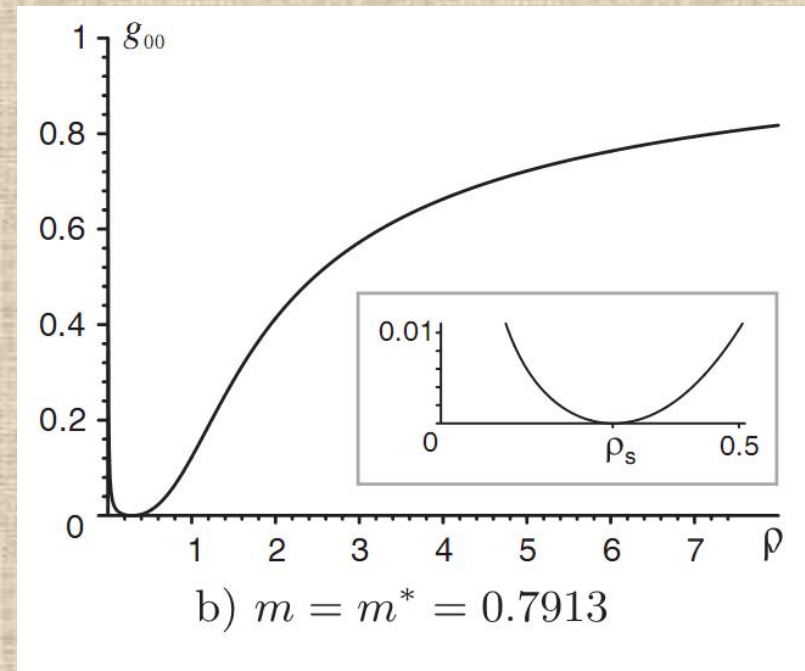
$$m_0 = \frac{1}{2} J(0) = 0.7924$$



Second specific value of the asymptotic mass



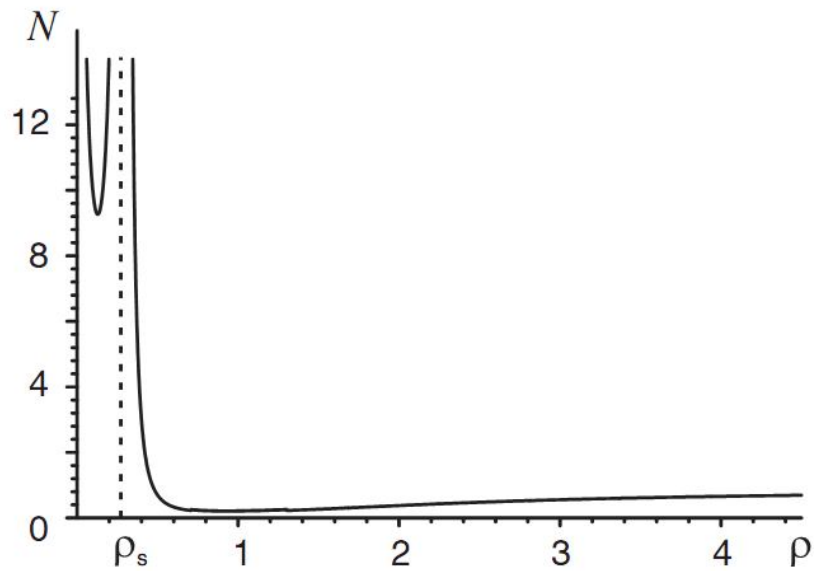
**$N(r)$ is positive
for this asymptotic mass**



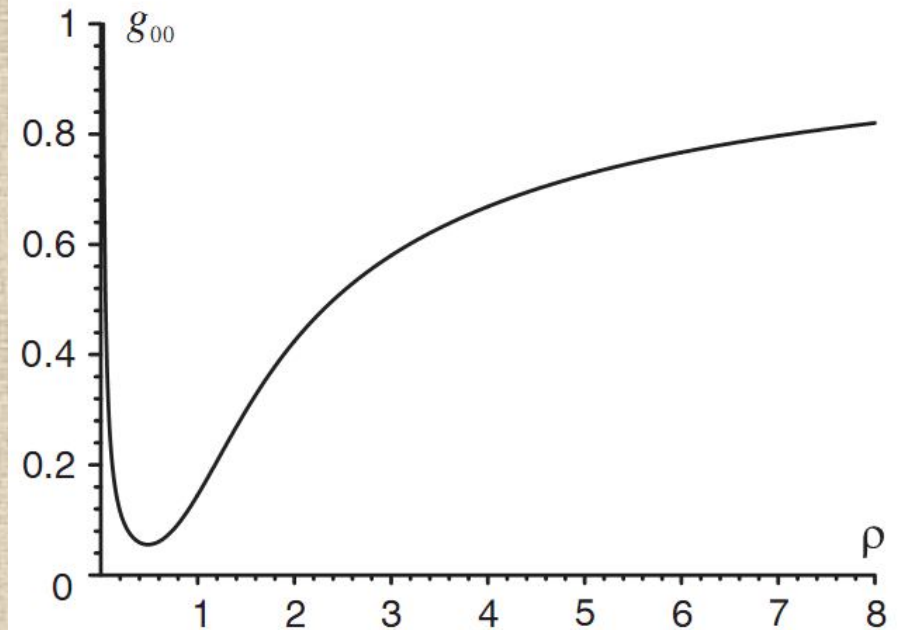
**This metric function is non-negative
and touches the horizontal axis at**

$$\rho = \rho_s = 0.289$$

Asymptotic masses as guiding parameters of the model:
I. No horizons. Central singularity. Infinite barrier.

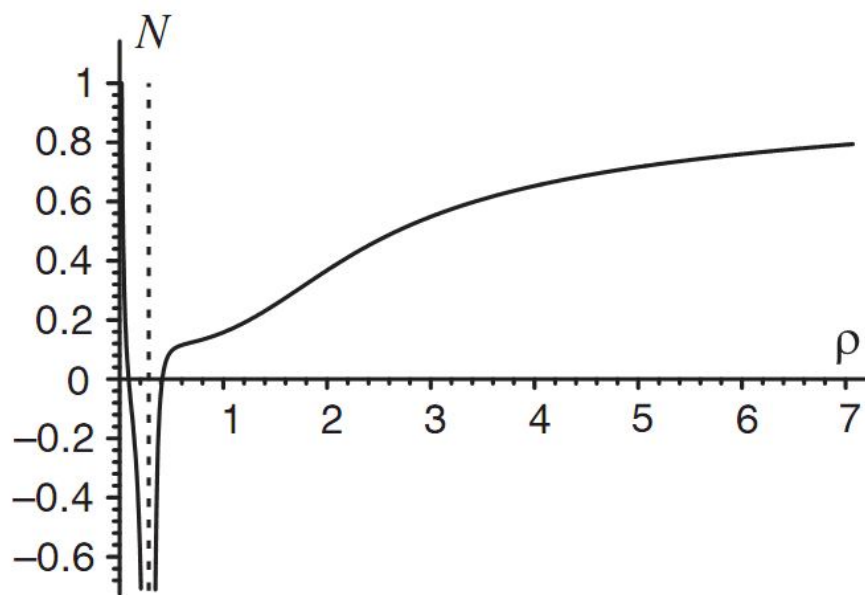


a) $0 < m < m^*$

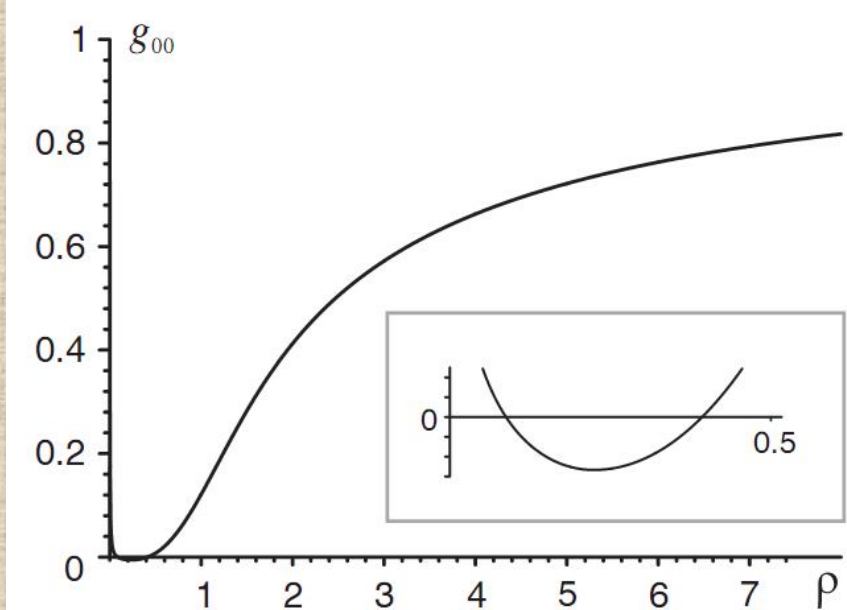


a) $0 < m < m^*$

Asymptotic masses as guiding parameters of the model:
II. Two horizons. Central singularity. Infinite well. Infinite barrier.



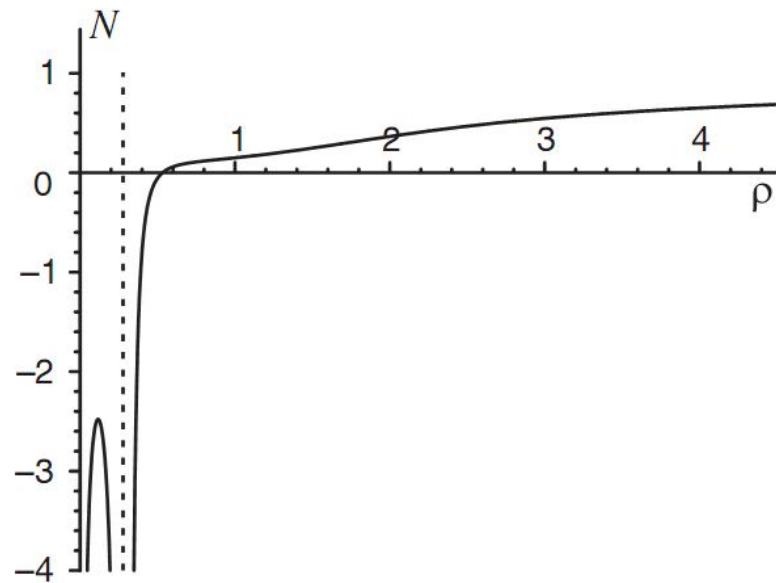
c) $m^* < m < m_0$



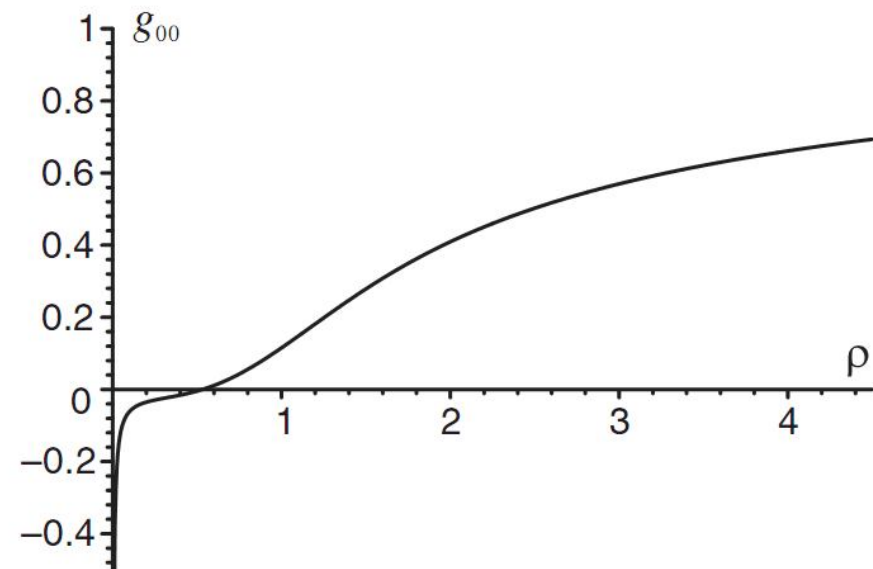
c) $m^* < m < m_0$

Asymptotic masses as guiding parameters of the model:

III. One horizon. Central singularity. Infinite well.



e) $m > m_0$



e) $m > m_0$

Magnetic analog: non-minimal Wu-Yang monopole with regular gravitational field and magnetic field singular in the center

A.B. Balakin and A.E. Zayats. Phys. Lett. B, 2007.

$$q_1 = -q < 0$$

$$q_2 = 4q$$

$$q_3 = -6q$$

$$B = \frac{\nu}{\mathcal{G}r^2}$$

$$\sigma(r) = 1$$

$$N(x) = 1 + \frac{x^2(1 - mx)}{1 + x^4}$$

$$N(\infty) = 1$$

$$N(0) = 1$$

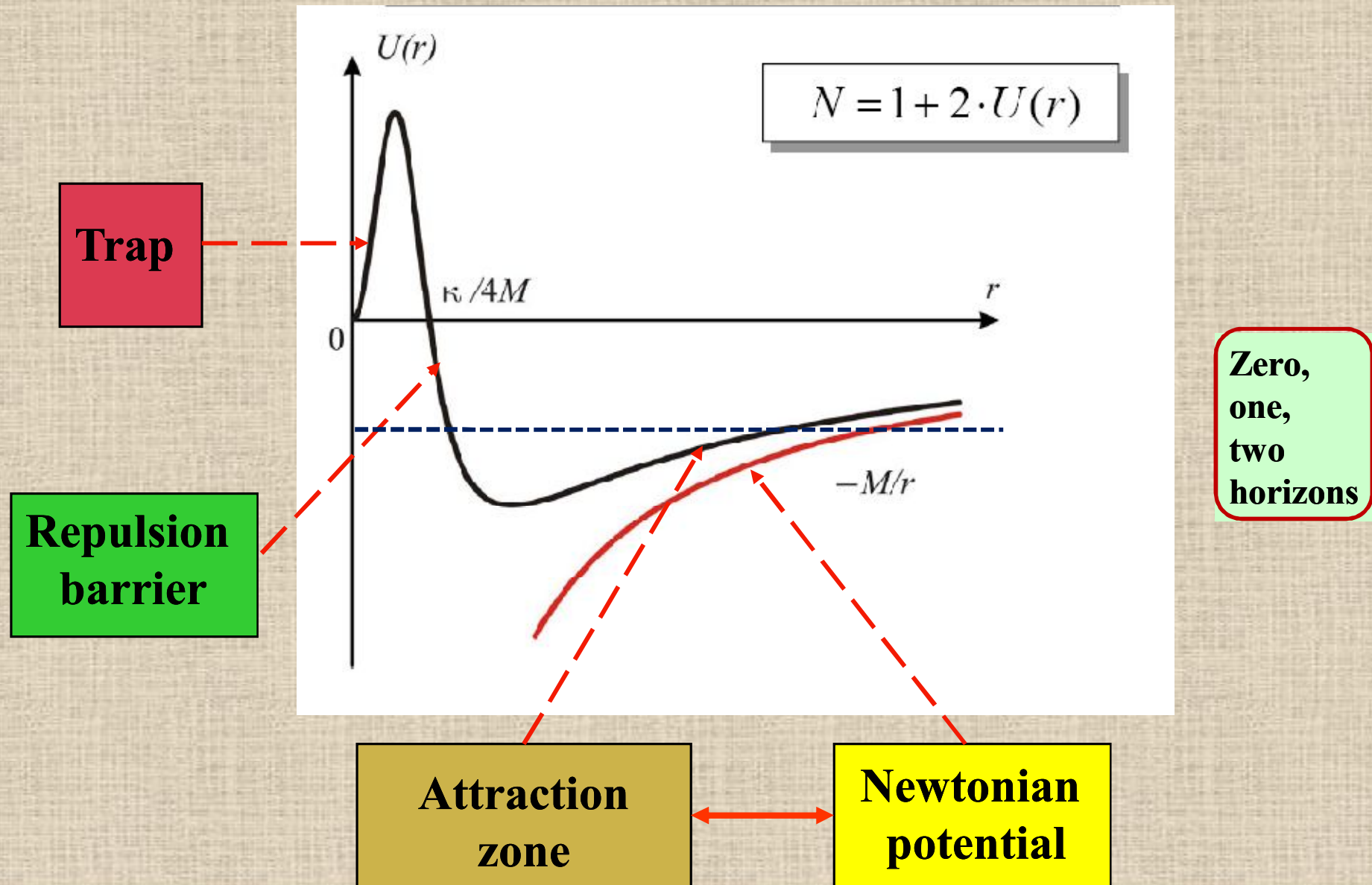
$$\kappa = 8\pi\nu^2/\mathcal{G}^2$$

$$r_Q^2 = \frac{\kappa}{2}$$

$$m = \frac{2M}{r_Q}$$

$$2q = r_Q^2$$

Effective gravitational potential near the non-minimal regular Wu-Yang monopole



Summary

- We considered three special non-minimal models, for which the Reissner-Nordstrom and non-minimal radii coincide; the obtained **exact** solutions for the electric, magnetic and gravitational fields are **analytical**.
- For the first model the electric field and both metric functions are **regular in the center**, but curvature invariants are singular.
- For the second model the electric field is again **regular in the center**, but one of the metric functions is singular.
- For the third model the metric functions and curvature invariants are **regular in the center**, but the magnetic field is singular.

Is there no perfection in this world?

We intend to continue to search for **completely regular exact solutions** using the model of non-minimal **dyons** and the model of **axion electrostatics** !!!

THANK YOU FOR THE ATTENTION !